

Filament Formation by Collision-induced Magnetic Reconnection (CMR)

THE ASTROPHYSICAL JOURNAL, 906:80 (28pp), 2021 January 10

© 2021. The American Astronomical Society. All rights reserved.

<https://doi.org/10.3847/1538-4357/abc687>



CrossMark

The CARMA-NRO Orion Survey: Filament Formation via Collision-induced Magnetic Reconnection—the Stick in Orion A

Shuo Kong (孔朔)^{1,2} , Volker Ossenkopf-Okada³ , Héctor G. Arce² , John Bally⁴ , Álvaro Sánchez-Monge³ ,
Peregrine McGehee⁵ , Sümeyye Suri⁶ , Ralf S. Klessen^{7,8} , John M. Carpenter⁹ , Dariusz C. Lis¹⁰ , Fumitaka Nakamura¹¹ ,
Peter Schilke³ , Rowan J. Smith¹² , Steve Mairs¹³ , Alyssa Goodman¹⁴ , and María José Maureira¹⁵ 

CMR exploration I – filament structure with synthetic observations

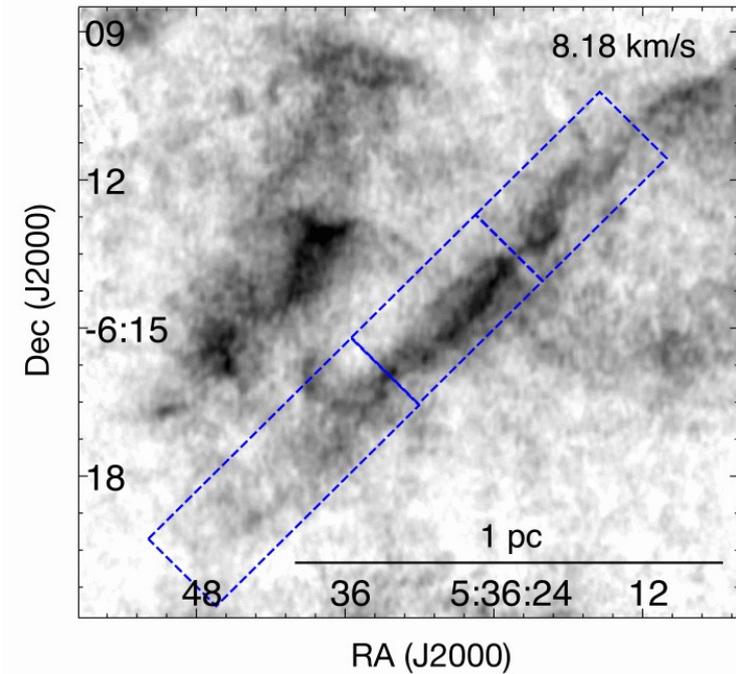
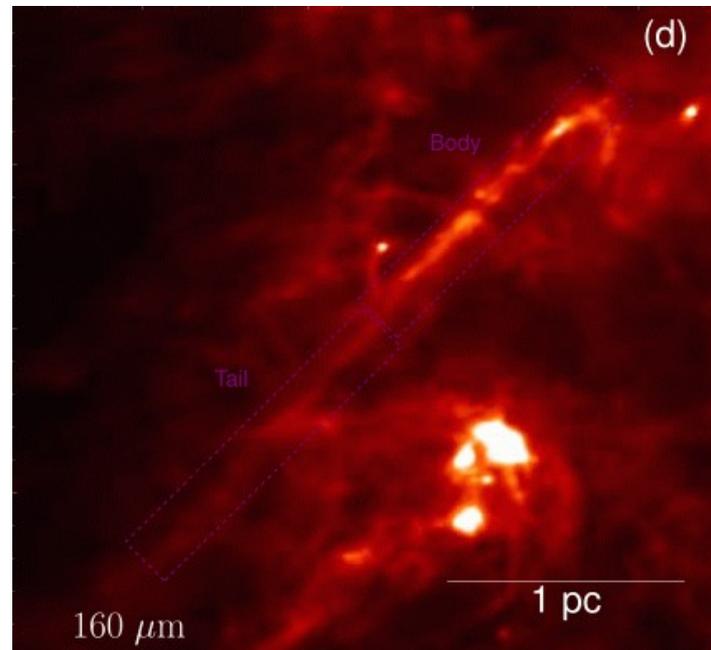
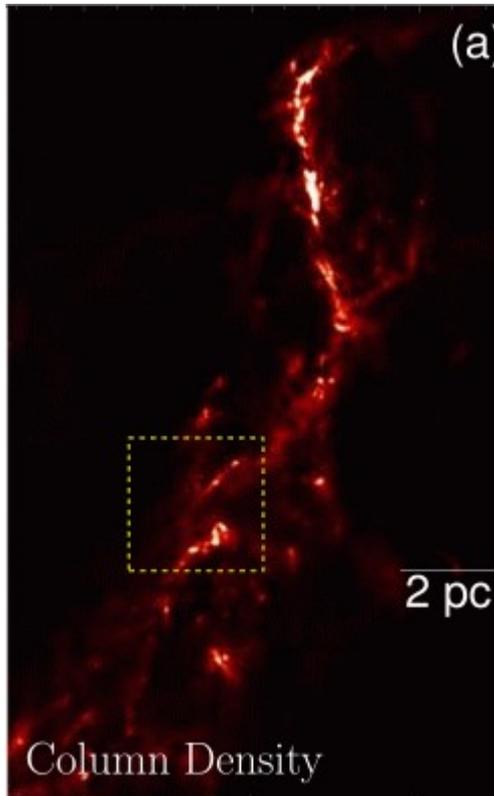
SHUO KONG (孔朔),¹ VOLKER OSSENKOPF-OKADA,² HÉCTOR G. ARCE,³ RALF S. KLESSEN,^{4,5} AND
DUO XU⁶

CMR exploration II – filament identification with machine learning

DUO XU,¹ SHUO KONG,² AVICHAL KAUL,² HÉCTOR G. ARCE,³ AND VOLKER OSSENKOPF-OKADA⁴

Observational motivation

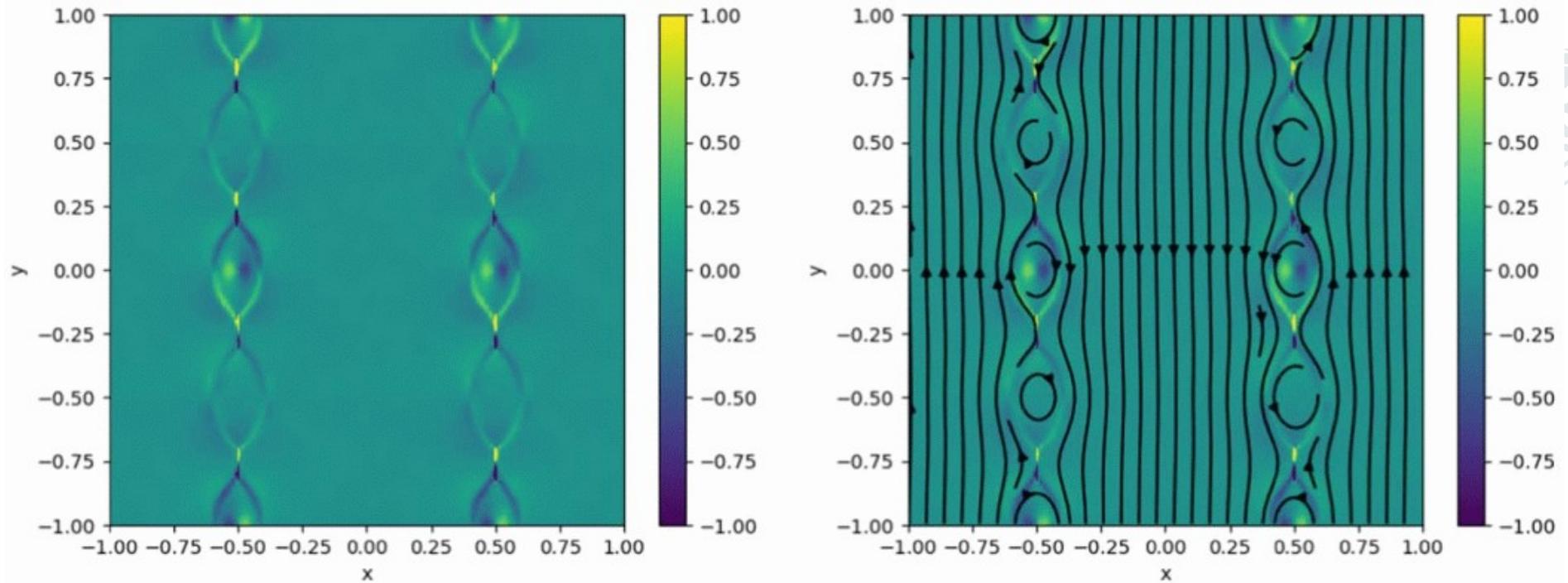
- Large-scale Orion mapping
 - prominent very straight filament close to L1630



- Breaks up into rings and forks, in particular in velocity slices

Association

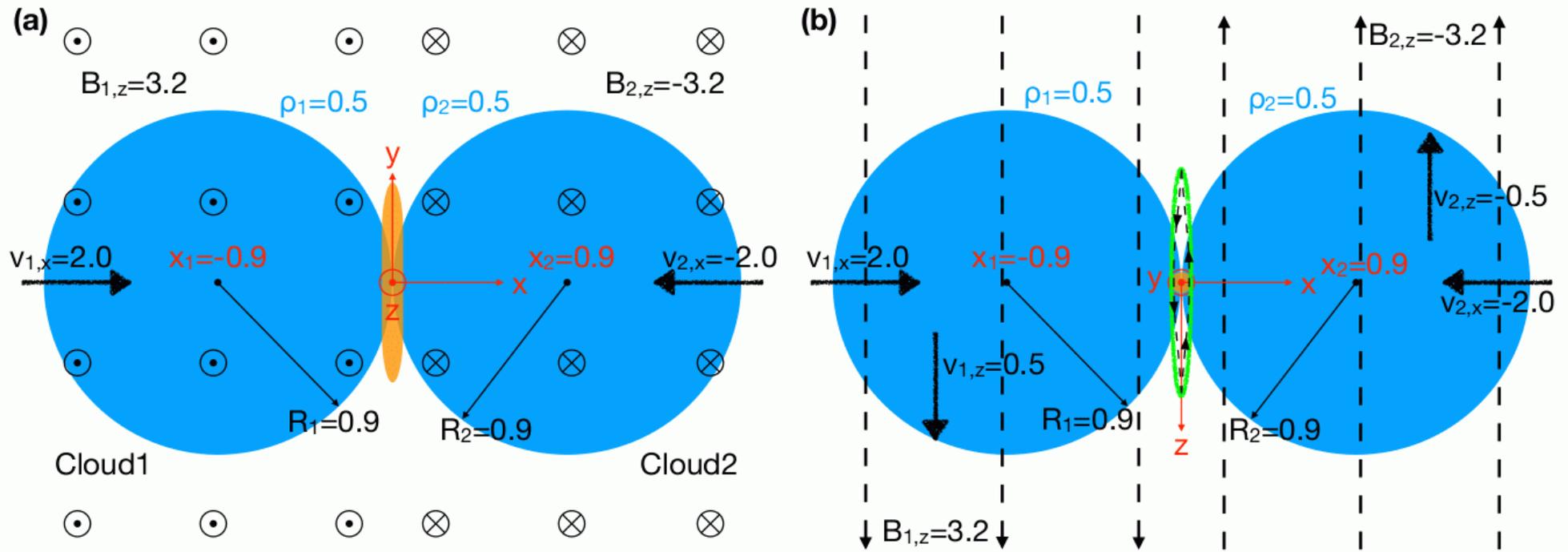
- Magnetic reconnection
 - Structure formation at the boundary of opposite magnetic fields



- Velocity pattern in MR

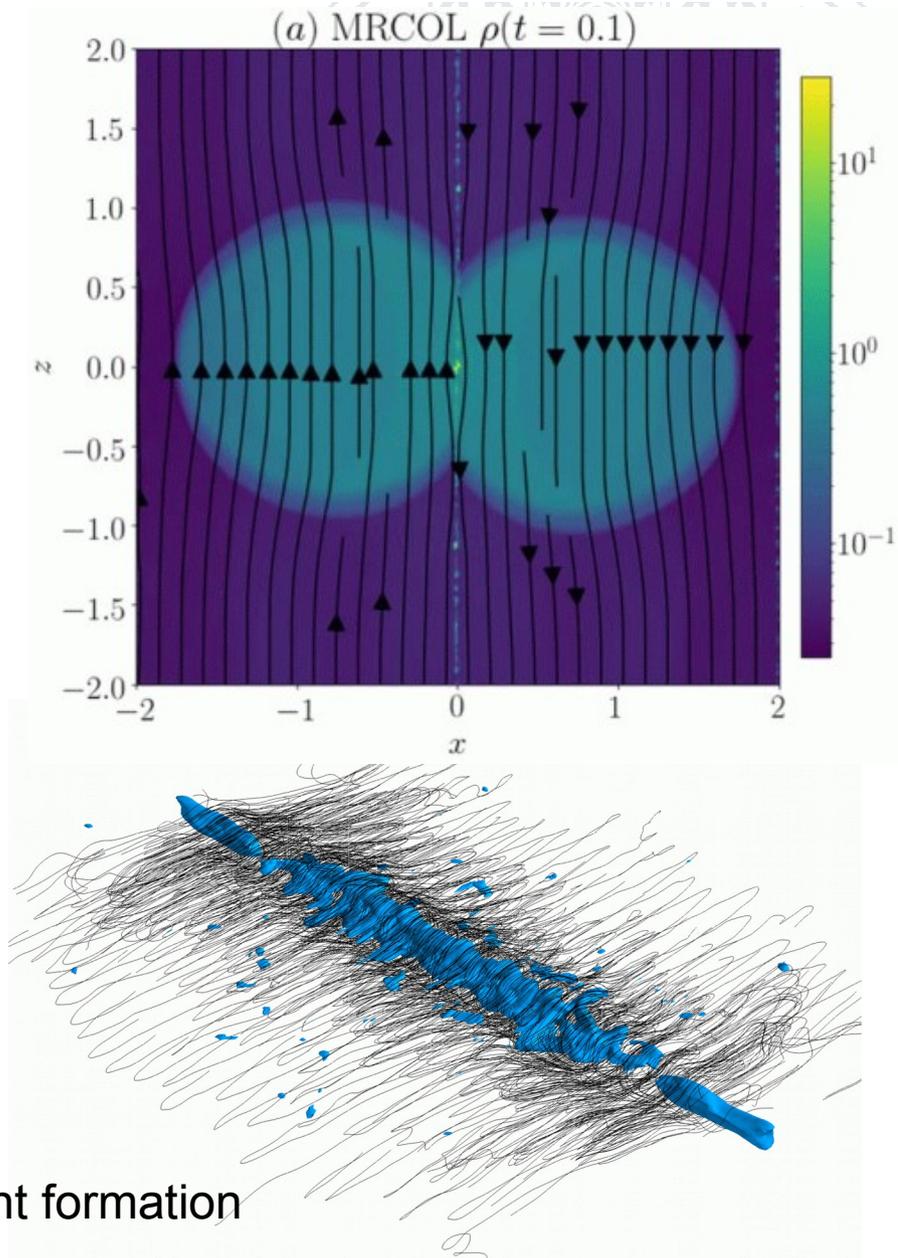
Fundamental model

- Collision-induced Magnetic Reconnection (CMR)
 - Filament forms perpendicular to velocity and B field



Simulation

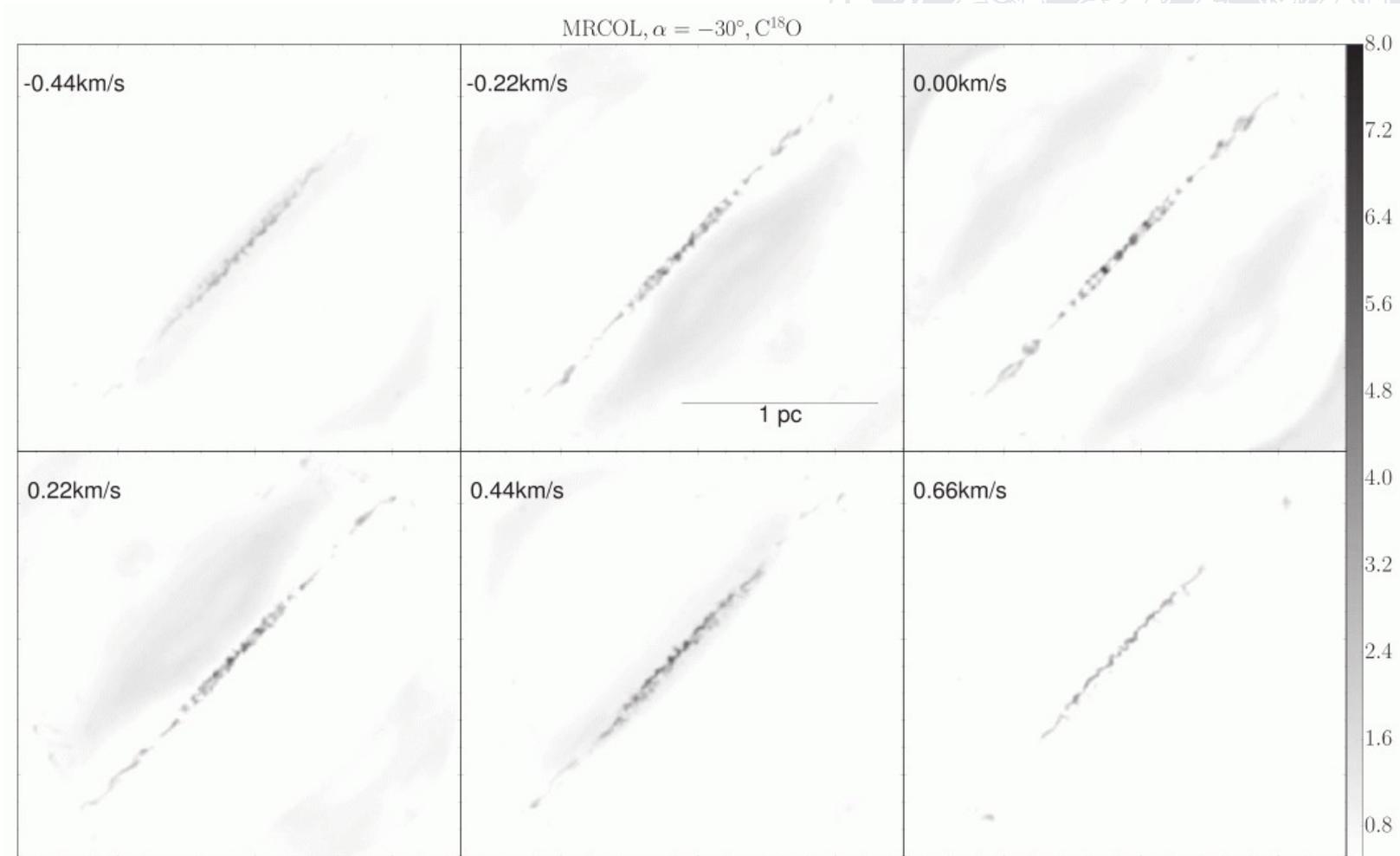
- ATHENA++ (Stone et al. 2020)
 - Resistive MHD
 - isothermal
 - uniform Cartesian grid
 - 512^3 voxels: $0.0078\text{pc} = 1600\text{AU} \rightarrow 4\text{pc}$ box
 - Default parameters: $n_{\text{H}}=840\text{ cm}^{-2}$, $T=15\text{K}$, $v=0.51\text{ km/s}$, $B=10\mu\text{G}$
 - Radiative transfer with RADMC-3D and Simline3D



Features reproduced

- Rings and forks in particular in velocity channels

- Process in Orion supported by Zeeman measurements



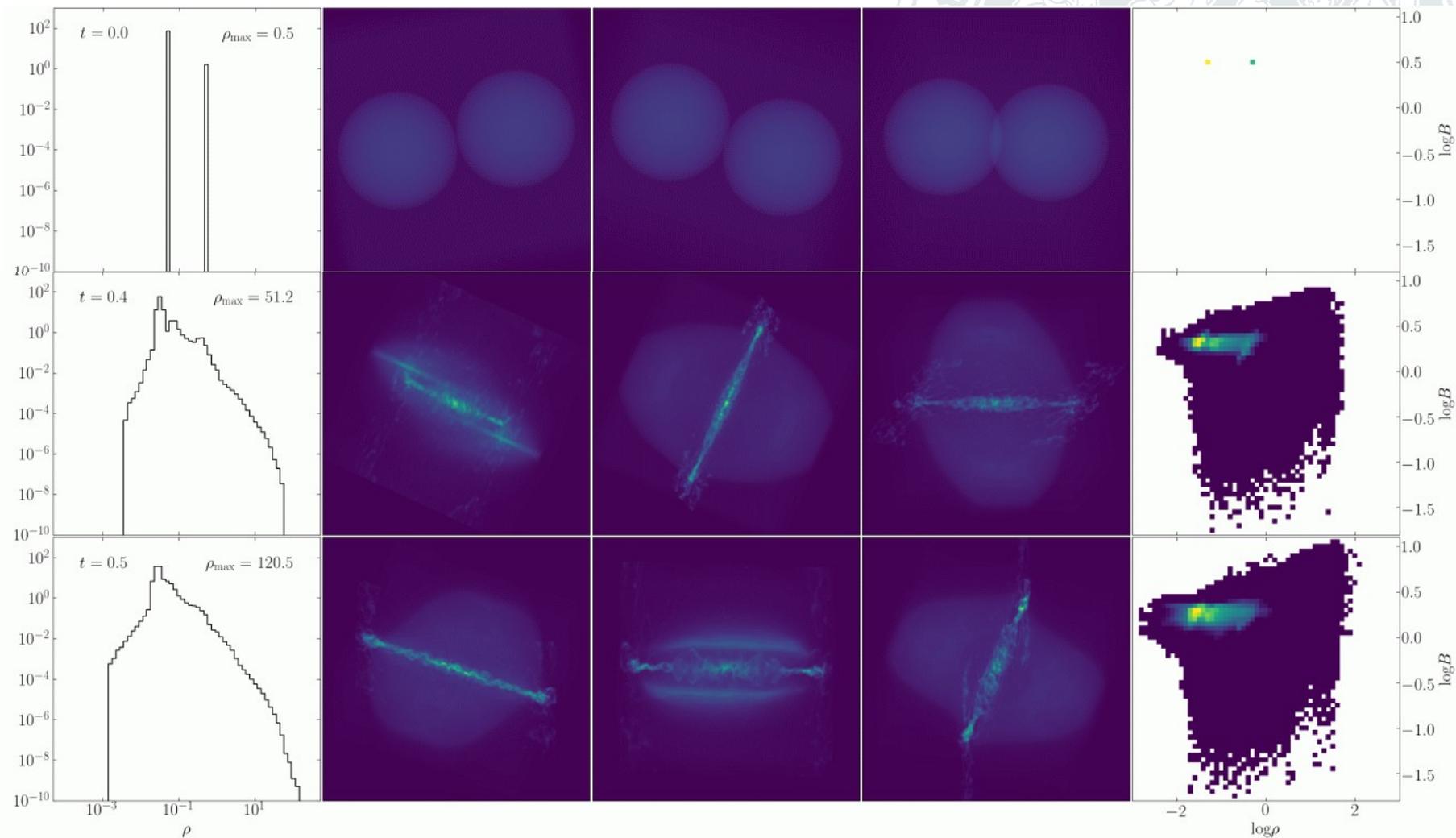
Systematics

- Variation of parameters, like collision angle – study 400 different viewing angles

Model	#	T	η	Cloud1					Cloud2				
				ρ_1	R_1	$v_{1,x}$	$v_{1,z}$	$B_{1,z}$	ρ_2	R_2	$v_{2,x}$	$v_{2,z}$	$B_{2,z}$
MRCOL (§1)	(0)	15	0.001	0.5	0.9	2.0	0.5	3.2	0.5	0.9	-2.0	-0.5	-3.2
η _L (§4.1)	(1)	15	<u>0.0001</u>	0.5	0.9	2.0	0.5	3.2	0.5	0.9	-2.0	-0.5	-3.2
η _H (§4.1)	(2)	15	<u>0.01</u>	0.5	0.9	2.0	0.5	3.2	0.5	0.9	-2.0	-0.5	-3.2
B _L (§4.2)	(3)	15	0.001	0.5	0.9	2.0	0.5	<u>1.6</u>	0.5	0.9	-2.0	-0.5	<u>-1.6</u>
B _H (§4.2)	(4)	15	0.001	0.5	0.9	2.0	0.5	<u>6.4</u>	0.5	0.9	-2.0	-0.5	<u>-6.4</u>
ρ_2 _L (§4.3)	(5)	15	0.001	0.5	0.9	2.0	0.5	3.2	<u>0.25</u>	0.9	-2.0	-0.5	-3.2
ρ_2 _H (§4.3)	(6)	15	0.001	0.5	0.9	2.0	0.5	3.2	<u>1.0</u>	0.9	-2.0	-0.5	-3.2
R_2 _L (§4.4)	(7)	15	0.001	0.5	0.9	2.0	0.5	3.2	0.5	<u>0.45</u>	-2.0	-0.5	-3.2
R_2 _H (§4.4)	(8)	15	0.001	0.5	0.9	2.0	0.5	3.2	0.5	<u>1.8</u>	-2.0	-0.5	-3.2
T _L (§4.5)	(9)	<u>10</u>	0.001	0.5	0.9	2.0	0.5	3.2	0.5	0.9	-2.0	-0.5	-3.2
T _H (§4.5)	(10)	<u>30</u>	0.001	0.5	0.9	2.0	0.5	3.2	0.5	0.9	-2.0	-0.5	-3.2
v_x _L (§4.6)	(11)	15	0.001	0.5	0.9	<u>1.0</u>	0.5	3.2	0.5	0.9	<u>-1.0</u>	-0.5	-3.2
v_x _H (§4.6)	(12)	15	0.001	0.5	0.9	<u>4.0</u>	0.5	3.2	0.5	0.9	<u>-4.0</u>	-0.5	-3.2
v_z _L (§4.7)	(13)	15	0.001	0.5	0.9	2.0	<u>0.25</u>	3.2	0.5	0.9	-2.0	<u>-0.25</u>	-3.2
v_z _H (§4.7)	(14)	15	0.001	0.5	0.9	2.0	<u>1.0</u>	3.2	0.5	0.9	-2.0	<u>-1.0</u>	-3.2

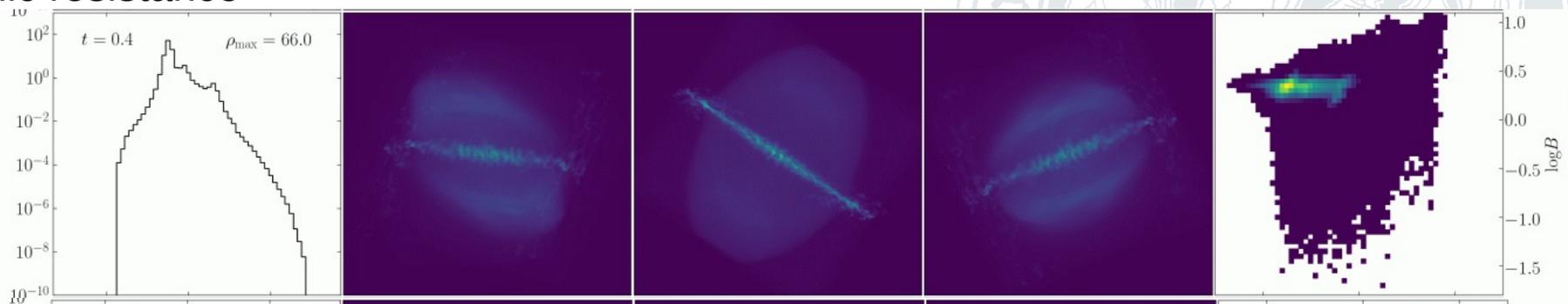
Examples

- Fiducial case (#0)

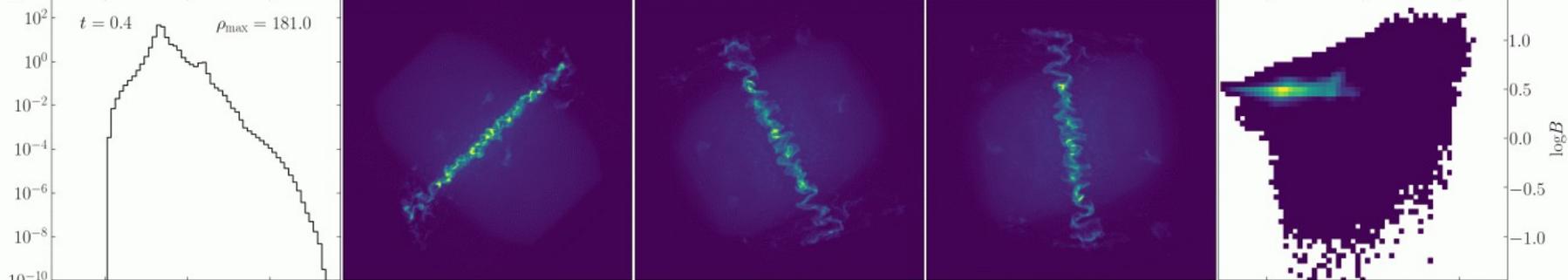


Examples

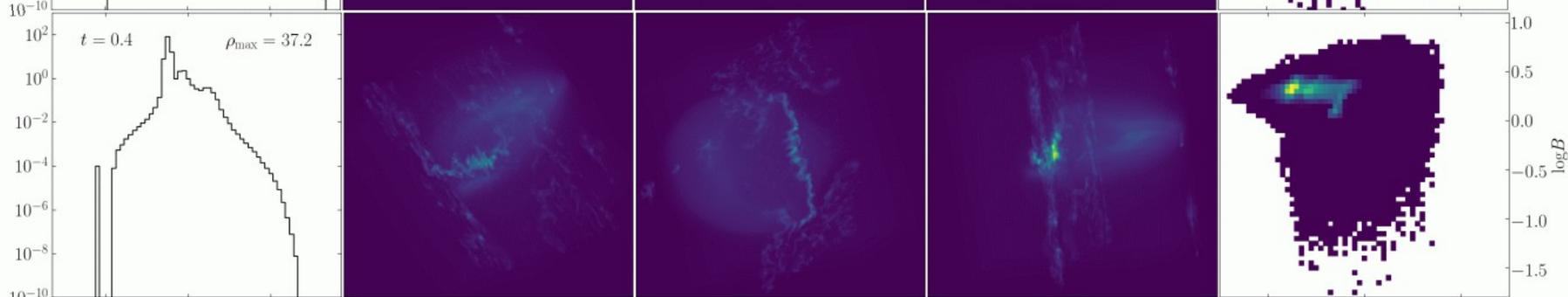
- Low Ohmic resistance (#1)



- High B field (#4)

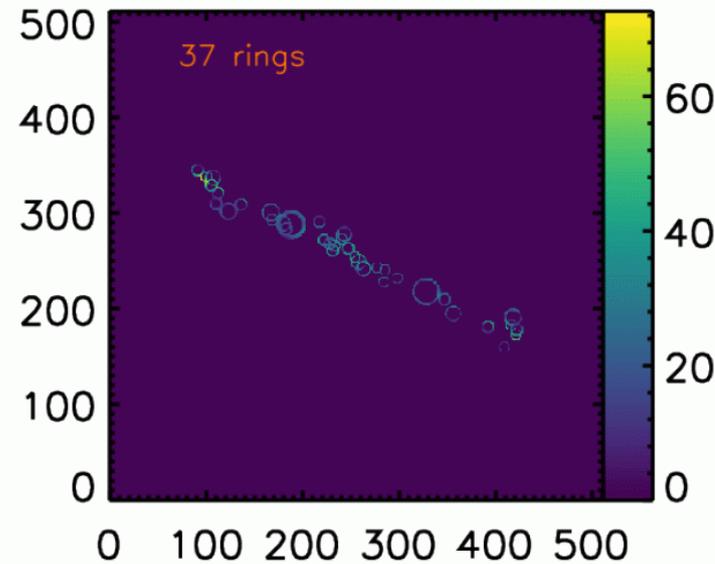
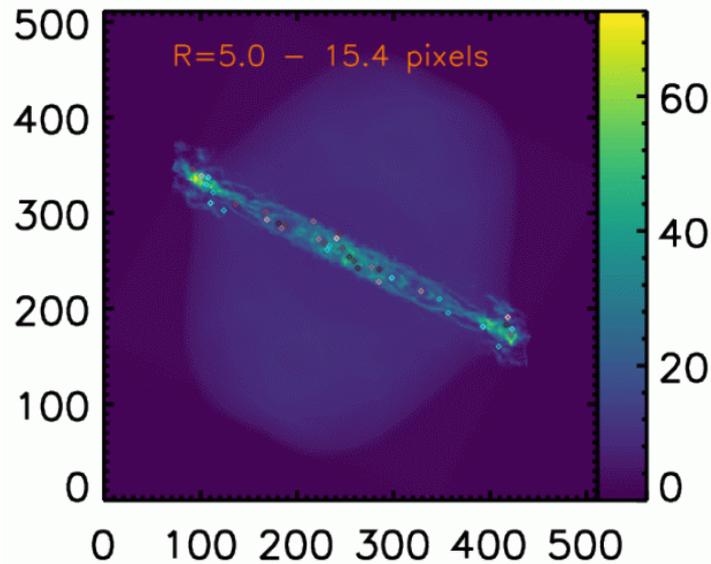
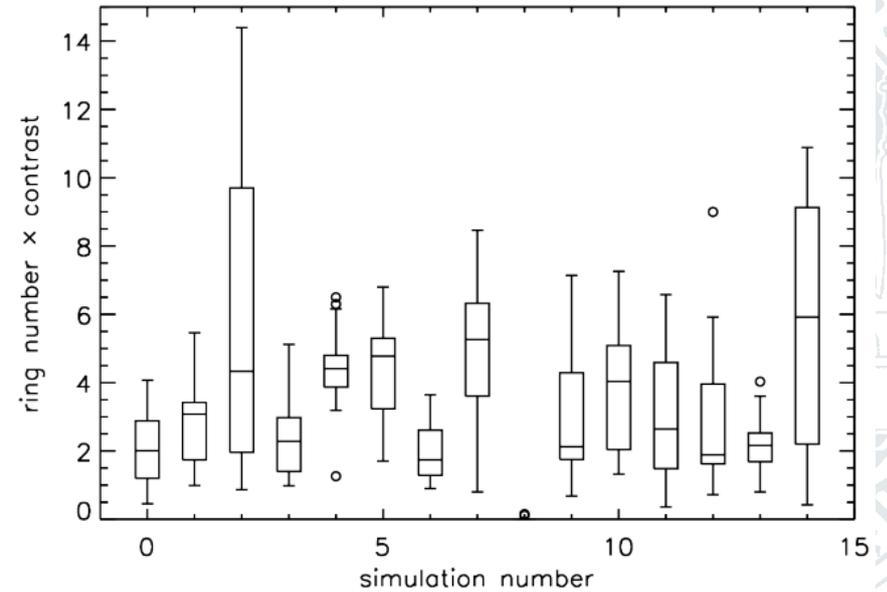


- 1/8 mass ratio (#7)



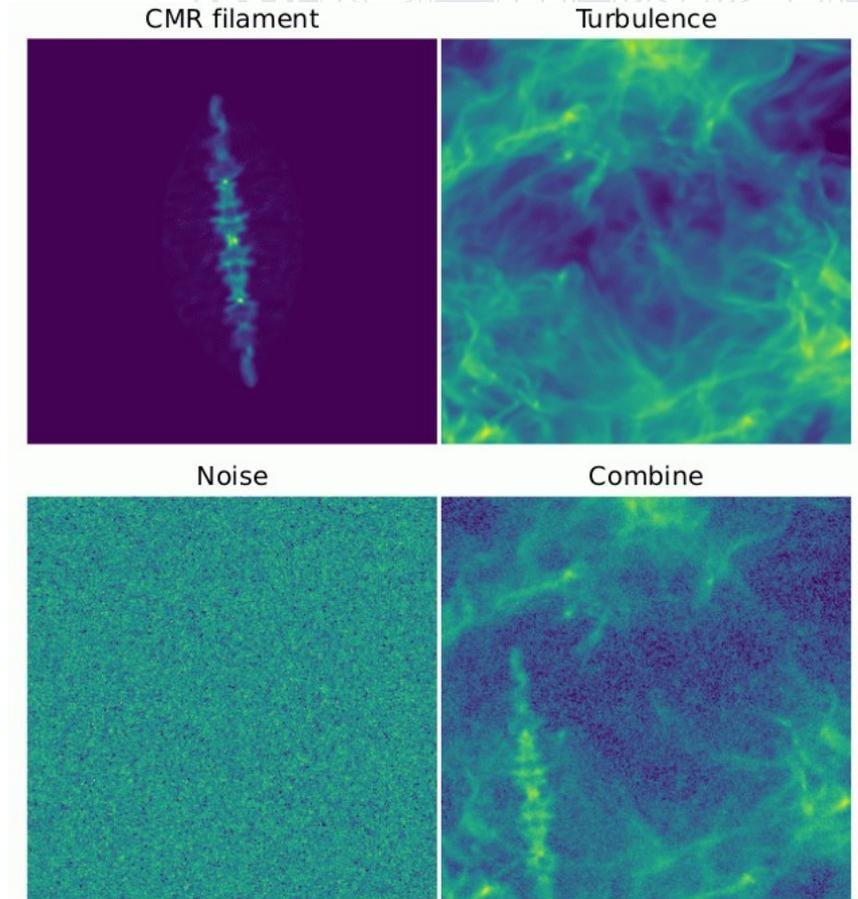
Systematics

- Analysis of “ringiness”
 - Wavelet-based ring-finder



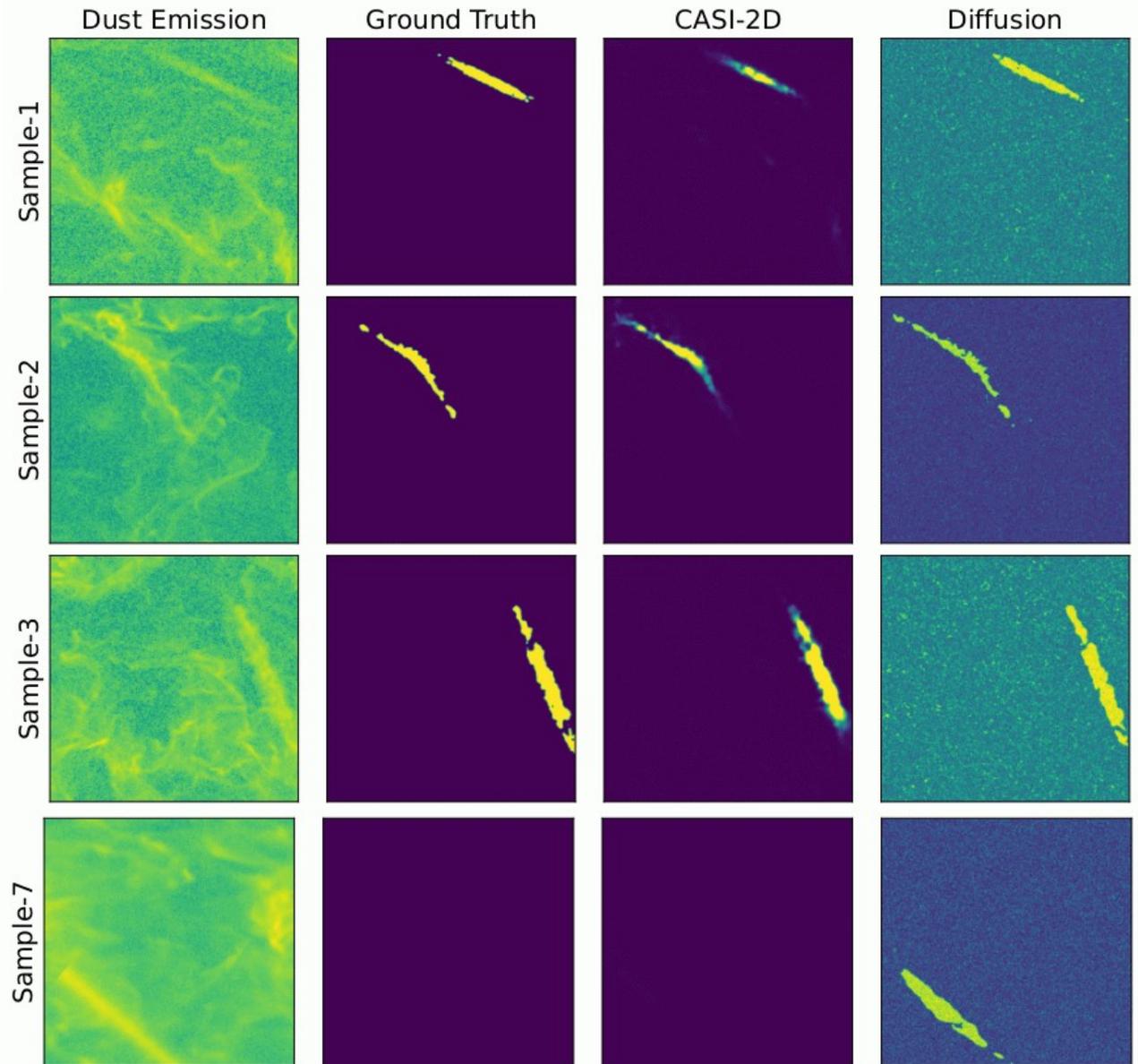
Better systematics

- Look for CMR filaments in observational data
 - Train CNN to find the filaments
 - 2 approaches
 - CASI-2D
 - Denoising Diffusion Probabilistic Model (DDPM)
 - Training data:
 - 14 simulations with 400 viewing angles
 - Negative set: MHD simulations of random turbulent boxes



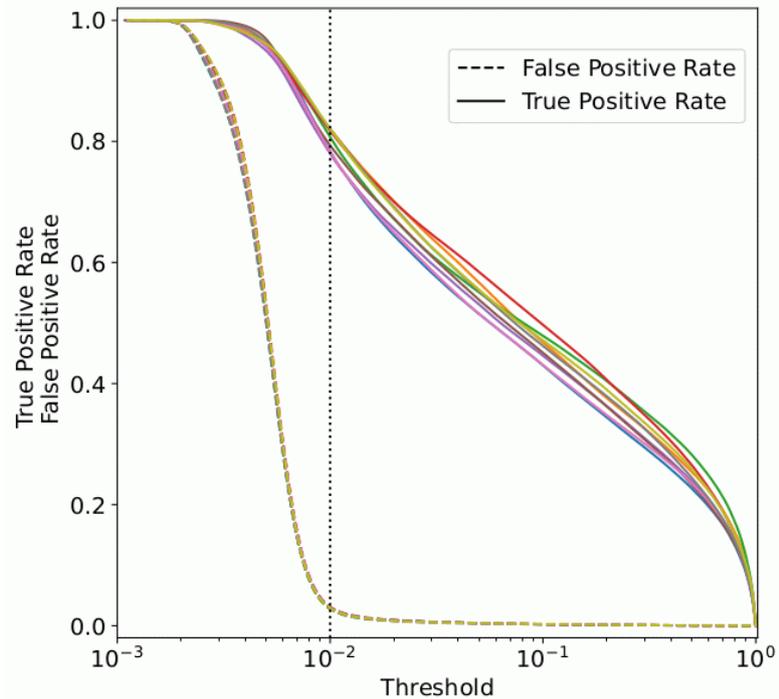
Results

- Network output giving the detection level for CMR filament membership of individual pixels
 - false detections in individual network output
 - often at map edges
 - Combining both networks eliminates most false detections

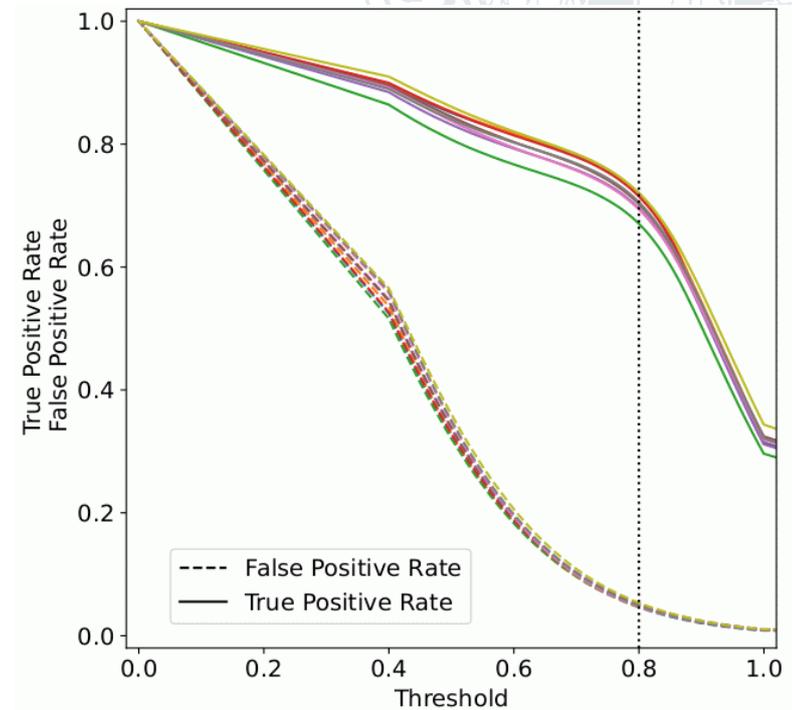


Optimum threshold

- Receiver Optimum Characteristics (ROC) curves
 - Different thresholds for the two networks: force false detections < 6%



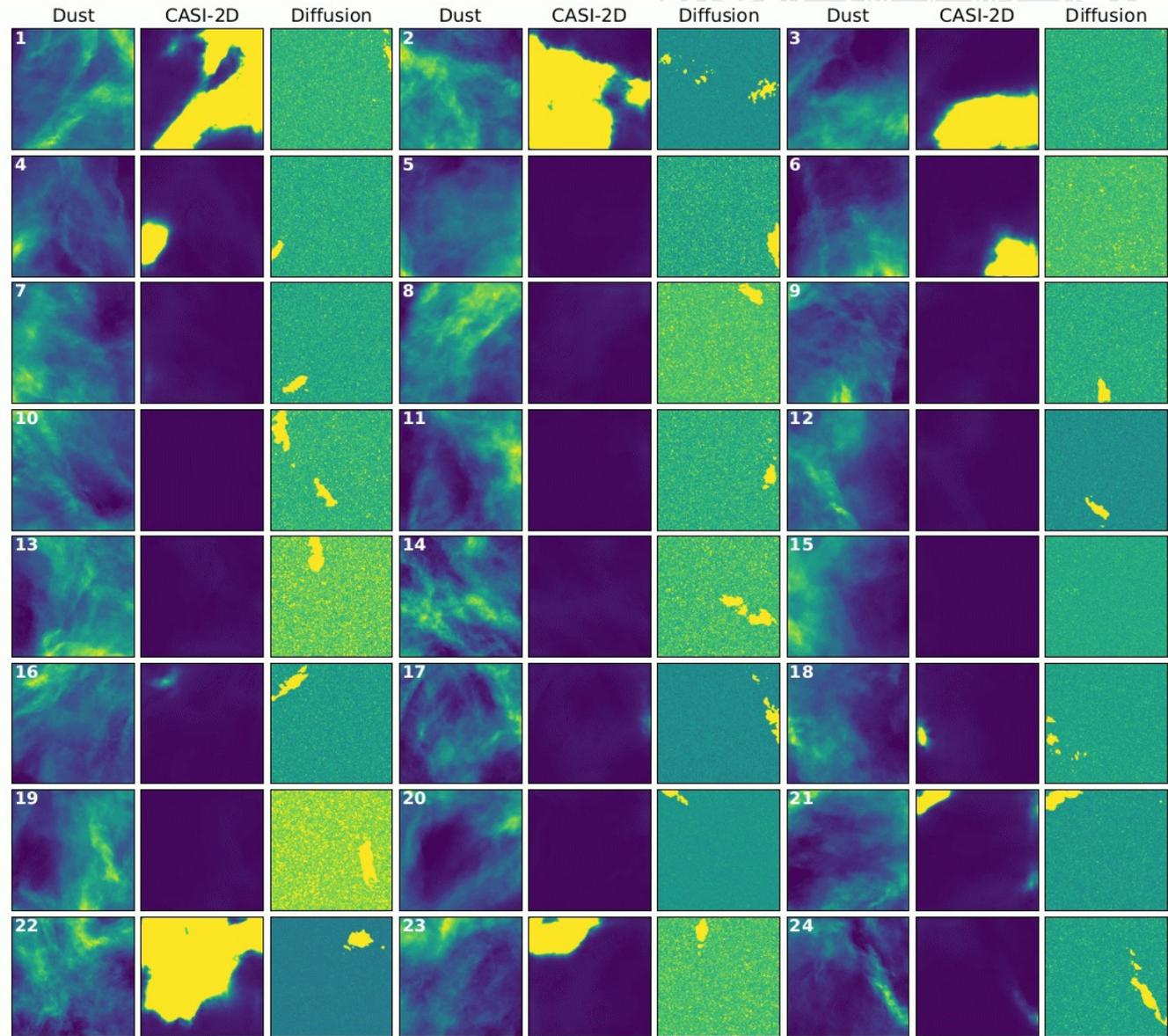
CASI-2D



diffusion model

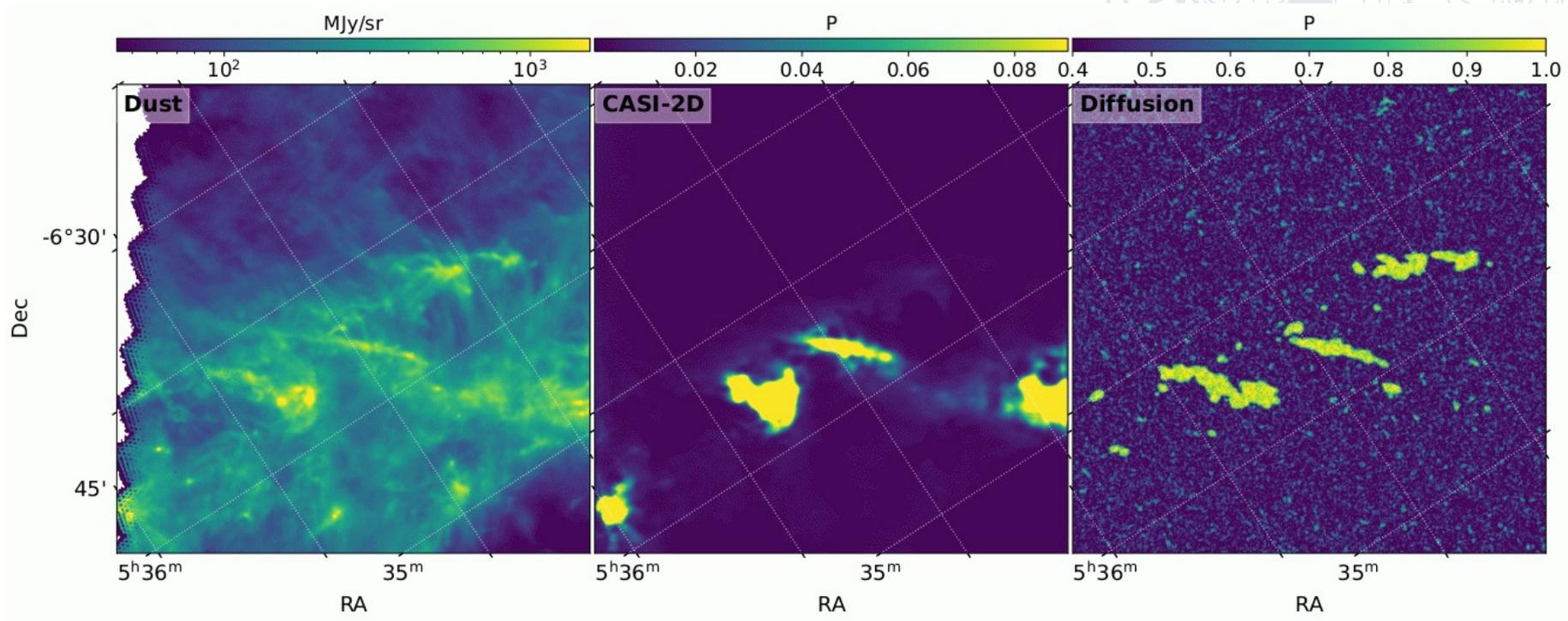
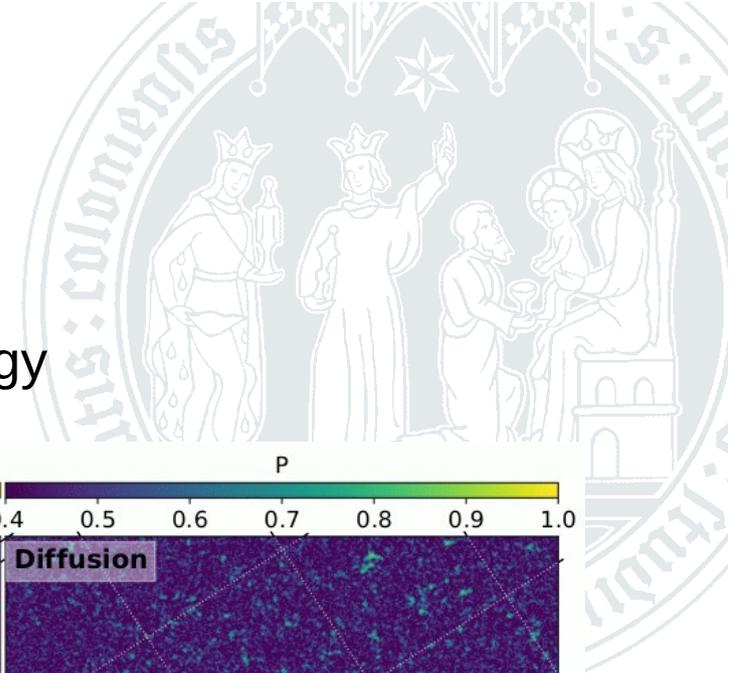
Fake identifications

- Test with HD simulations (Federrath 2021)
 - No CMR
 - Combining both networks eliminates most false detections
 - Problems at edges



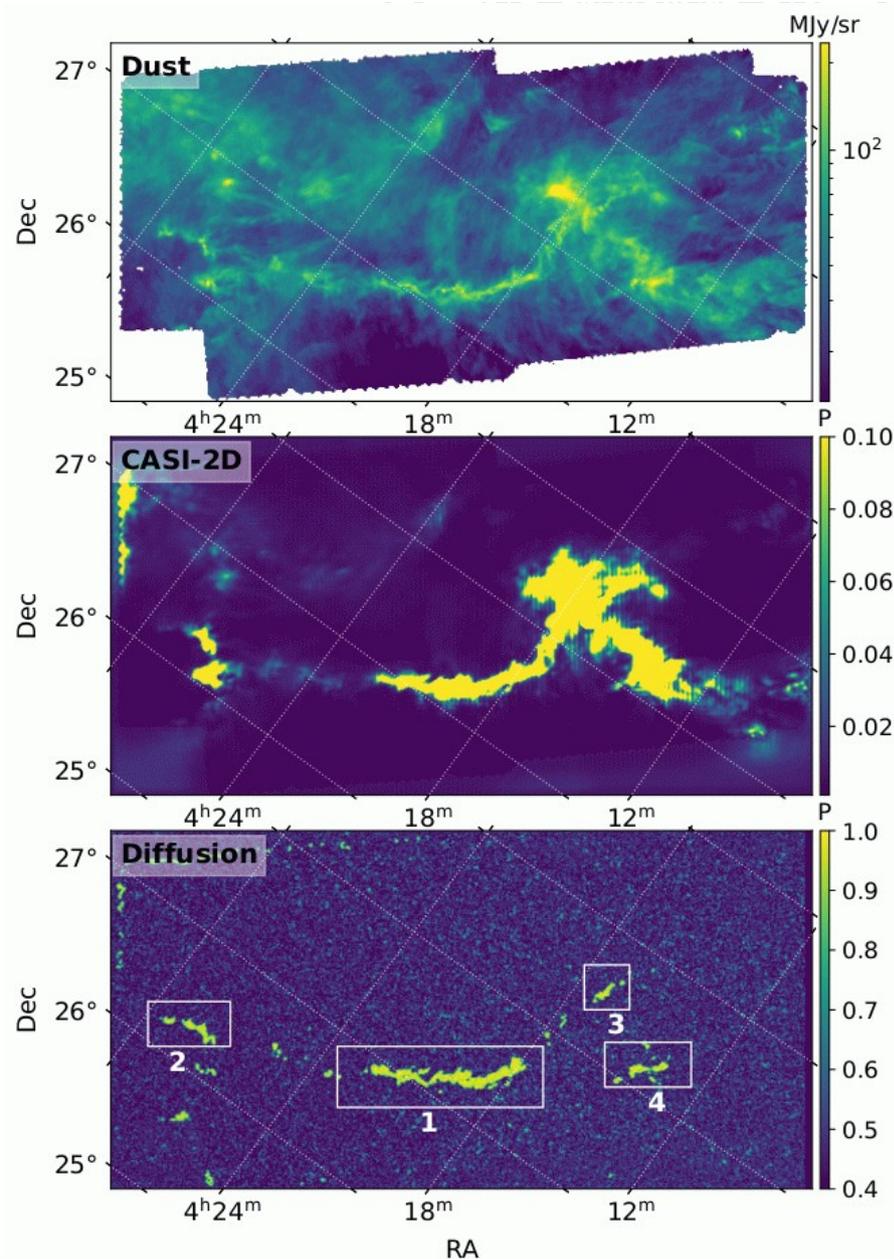
Known case

- Stick well detected
 - Other peaks can be excluded based on morphology



Application: Taurus

- B213 as candidate for high probability CMR filament
 - Others easily excluded based on morphology



Conclusions

- Two approaches (CASI-2D and diffusion model) need to be combined to obtain reasonable reliability
- Identification of candidates for CMR filaments
 - Confirmation needs measurement of magnetic field direction

